## Geodesy 1 (GED203)

## Lecture No: 5

# GEODETIC POSITION COMPUTATION 

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## Review of Previous Lecture

REVIEW OF PLANE TRIANGLE FORMULAS
PLANE TRIANGLES VERSUS SPHERICAL/GEODETIC TRIANGLES

## SPHERICAL TRIANGLE

SOLUTION OF SPHERICAL TRIANGLE

NAPIER'S RULE FOR RIGHT-ANGLE SPHERICAL TRIANGLES

SPHERICAL EXCESS

CALCULATION OF SPHERICAL EXCESS

## ELLIPSOIDAL EXCESS

## SUMMARY



## WHAT IS MEANT BY GEODETIC COMPUTATION?

## TYPES OF GEODETIC PROBLEMS.

## METHODS OF COMPUTATION

## ACCURACY OF THE EXISTING METHODS

## SUMMARY

## Expected Learning OUtcomes

- Gain a clear understanding of what is meant by the geodetic problem.
- Explore the various types of geodetic problems encountered in geodesy.
- Acquire knowledge about the different computation methods used to solve geodetic problems.
- Learn about the accuracy of existing computation methods in geodesy.
- Apply the knowledge of geodetic problems and computation methods to real-world scenarios.
- Develop critical thinking skills to analyze and solve geodetic problems effectively.
- Collaborate with peers to discuss and solve geodetic problems, share insights, and explore new approaches.


## What is Geodetic Problem?

- The main problem of geodesy is the computation of the geodetic coordinates ( $\varphi, \lambda, \mathrm{h}$ ) of different points on the reference ellipsoid.
- This computation necessitates measurements of angles ( $\mathrm{H} \& \mathrm{~V}$ ), azimuths, and distances.
- Using such measurements at specific locations, the coordinates of the remaining points in the network can be computed based on triangulation.



## Types of Geodetic Problems

- Two cases of calculations exist:
a) Direct Geodetic Problem: if the coordinates of a starting point, a distance and the forward azimuth to a second point are given and it is required to find the coordinates of the second point as well as the backward azimuth.
b) Inverse Geodetic Problem: if the coordinates of the two points are known and it required to find the length of the line and its forward and back azimuth.
- The solution of either of these problems is a solution of the ellipsoidal polar triangle.



## Types of Geodetic Problems

- Direct Problem:

$$
\begin{gathered}
\varphi_{2}=f_{1}\left(\varphi_{1}, \lambda_{1}, \alpha_{12}, S\right) \\
\lambda_{2}=f_{2}\left(\varphi_{1}, \lambda_{1}, \alpha_{12}, S\right) \\
\alpha_{21}=f_{3}\left(\varphi_{1}, \lambda_{1}, \alpha_{12}, S\right)
\end{gathered}
$$

- Inverse Problem:

$$
\begin{gathered}
S=f_{4}\left(\varphi_{1}, \lambda_{1}, \varphi_{2}, \lambda_{2}\right) \\
\alpha_{12}=f_{5}\left(\varphi_{1}, \lambda_{1}, \varphi_{2}, \lambda_{2}\right) \\
\alpha_{21}=f_{6}\left(\varphi_{1}, \lambda_{1}, \varphi_{2}, \lambda_{2}\right)
\end{gathered}
$$



## How to Solve A Geodetic Problem

- On a sphere the solutions to both problems are (simple) exercises in spherical trigonometry. On an ellipsoid the computation is much more involved.
- Work on ellipsoidal solutions was carried out by for example Legendre, Bessel, Gauss, Laplace, Helmert and many others after them.
- There are many solutions for these problems.
- The solutions are classified w.r.t the distance connecting the two points and the type of this distance (Normal Section or Geodesic).


## Methods of Computations

- The most common methods in literature are:
I. Clarke's formula for long lines
II. Clarke's formula for medium and short lines
III. Puissant formula and modifications of them
IV. Lilly's formula for long lines
V. Mid-latitude formula
VI. Bowring formula for geodesic lines up to 150 km
VII. The chord method
VIII. Bessel's formula
IX. The Chord Method
- These formulas are deduced use on a computational reference ellipsoid for first and secondorder accuracies.
- For the case of adopting a sphere as a reference surface for geodetic computations, the simplification of these formula will yield results that are rigorous enough for third-order accuracies.


## Methods of Computations - Mid-Latitude Formula

## - Mid-latitude formula

It is first published in English in 1861. These formulas rely on a spherical approximation of the earth and should only be used for points separated by less than 40 km at latitudes less than $80^{\circ}$. The formula is simple to use and rigorous enough for second order accuracies.

The direct problem solution is an iterative solution for $d_{\varphi}$.
The inverse problem is computed without iteration since $\varphi_{m}$ is immediately available.

The accuracy of Gauss mid-latitude formulas is about 1 ppm for lines up to 100 km in length.

## Methods of Computations - Mid-Latitude Formula

- Mid-latitude formula (Direct Problem)

Assume

$$
\begin{align*}
\varphi_{m} & =\varphi_{1}, \Delta \alpha=0 \\
\Delta \lambda & =\frac{S \sin \left(\alpha_{12}+\Delta \alpha / 2\right)}{\left.N_{m} \cos \varphi_{m}\right)} \tag{1}
\end{align*}
$$

$$
\begin{equation*}
\Delta \varphi=\frac{S \cos \left(\alpha_{12}+\frac{\Delta \alpha}{2}\right)}{\left.M_{m} \cos \frac{\Delta \lambda}{2}\right)} \tag{2}
\end{equation*}
$$

$$
\begin{equation*}
\Delta \alpha=\Delta \lambda \times \sin \varphi_{m} \times \sec \left(\frac{\Delta \varphi}{2}\right)+\left(\frac{1}{12}\right) \sin \varphi_{m}\left(\cos ^{2} \varphi_{m}\right) \cdot \Delta \lambda^{3} \tag{3}
\end{equation*}
$$

## Methods of Computations - Mid-Latitude Formula

## - Mid-latitude formula (Direct Problem)

$N_{m}$
The mean radius of curvature in the prime vertical direction
$M_{m}$ $\qquad$ The mean radius of curvature in the meridian direction

Then,
$\varphi_{2}=\varphi_{1}+\Delta \varphi$
$\lambda_{2}=\lambda_{1}+\Delta \lambda$
$\alpha_{21}=\alpha_{12}+\Delta \alpha \pm 180^{\circ}$
repeat the solution using $\varphi_{m}=\frac{\left(\varphi_{1}+\varphi_{m}\right)}{2}$ and the new value of $\Delta \alpha$ until the value of $\varphi$ remains constant.

## Methods of Computations - Mid-Latitude Formula

- Mid-latitude formula (Inverse Problem)
$\varphi_{m}=\frac{1}{2}\left(\varphi_{1}+\varphi_{2}\right)$
$\Delta \alpha=\Delta \lambda \times \sin \varphi_{m} \times \sec \left(\frac{\Delta \varphi}{2}\right)+\left(\frac{1}{12}\right) \sin \varphi_{m}\left(\cos ^{2} \varphi_{m}\right) \cdot \Delta \lambda^{3}$
$M_{1}=S_{1} \cos \left(\alpha_{12}+\frac{\Delta \alpha}{2}\right)=M_{m} \cdot \Delta \varphi^{\prime} \cdot \cos \left(\frac{\Delta \lambda}{2}\right)$
$N_{1}=S_{1} \sin \left(\alpha_{12}+\frac{\Delta \alpha}{2}\right)=N_{m} \cdot \Delta \lambda^{\prime} \cdot \cos \left(\varphi_{m}\right)$
$S_{1}=\sqrt{N_{1}{ }^{2}+M_{1}{ }^{2}}$
$\Delta \varphi^{\prime}=\Delta \varphi\left[\sin \left(\frac{\Delta \varphi}{2}\right) /\left(\frac{\Delta \varphi}{2}\right)\right]$
$\Delta \lambda^{\prime}=\Delta \lambda\left[\sin \left(\frac{\Delta \lambda}{2}\right) /\left(\frac{\Delta \lambda}{2}\right)\right]$


## Methods of Computations - Mid-Latitude Formula

- Mid-latitude formula (Inverse Problem)

Now compute the unknowns $S, a_{12}$ and $\alpha_{21}$,

$$
\begin{align*}
& S=S_{1}\left[\frac{\left(\frac{S_{1}}{2 N_{m}}\right)}{\sin \left(\frac{S_{1}}{2 N_{m}}\right)}\right]  \tag{11}\\
& \alpha_{12}=\tan ^{-1}\left(\frac{N_{1}}{M_{1}}\right)-\left(\frac{\Delta \alpha}{2}\right)  \tag{12}\\
& \alpha_{21}=\alpha_{12}+\Delta \alpha \pm 180 \tag{11}
\end{align*}
$$

## Methods of Computations - Puissant Formulas

- Puissant Formulas
- The equations were derived by Puissant in the 18 th century.
- It is a short line formula.
- Convenient to use for short lines (up to 100 km )
- Results are in third-order accuracies
- Correct to 1 ppm for lines up to 50 km and 40 ppm for 250 km
- More suitable for direct problem


## Methods of Computations - Bessel's Formulas

- Bessel's Formulas
- The derivation is based on the geodesic on the ellipsoid.
- The accuracy of Bessel formula is not limited by the separation between the two points nor by the location of the points on the earth.
- The derivation is based on parameters on a sphere and the parameters on ellipsoid.


## Methods of Computations - Bowring Formulas

- Bowring Formulas
- Published in 1981by Bowring for geodesic lines up to 150 km in length.
- The method uses a conformal projection of the ellipsoid on a sphere called the Gaussian projection of the second kind.
- The scale factor is taken one at the starting point of the line.
- The geodesic from the ellipsoid in then projected to the corresponding line on the sphere where spherical trigonometry could be applied.


## Accuracy of Direct Problem Solution Using Different Methods

|  | 0.00001 | 0.0001 | 0.001 | 0.01 |
| :--- | :---: | :---: | :---: | :---: |
| Legendre Series <br> (4 terms) | 30 | 40 | 80 | 100 |
| Legendre Series <br> (5 terms) | 60 | 90 | 100 | 200 |
| Puissant <br> (short, 6.51$)$ | 10 | 10 | 10 | 10 |
| Puissant <br> $(l$ long, 6.40$)$ | 10 | 20 | 40 | 80 |
| Bowring <br> Chord | 70 | 100 | 700 |  |

## Maximum Error in The Solution of The Inverse Problem for Various Length Lines

| Line Length | Gauss Mid Latitude |  | Bowring |  |
| :---: | :---: | :---: | :---: | :---: |
| km | Azimuth(") | Distance (mm) | Azimuth(") | Distance $(\mathrm{mm})$ |
| 50 | 0.0048 | 4 | 0.0003 | 0.1 |
| 100 | 0.020 | 33 | 0.0024 | 1.1 |
| 200 | 0.083 | 136 | 0.0049 | 9.7 |

## Important Resources

- https://pypi.org/project/geopy/
- https://geographiclib.sourceforge.io/
- https://pyproj4.github.io/pyproj/stable/
- https://pypi.org/project/PyGeodesy/


# LET'S SUMMARIZE 



## THANK YOU

End of Presentation

